

# A Double radiator configuration approach for the RICH detector

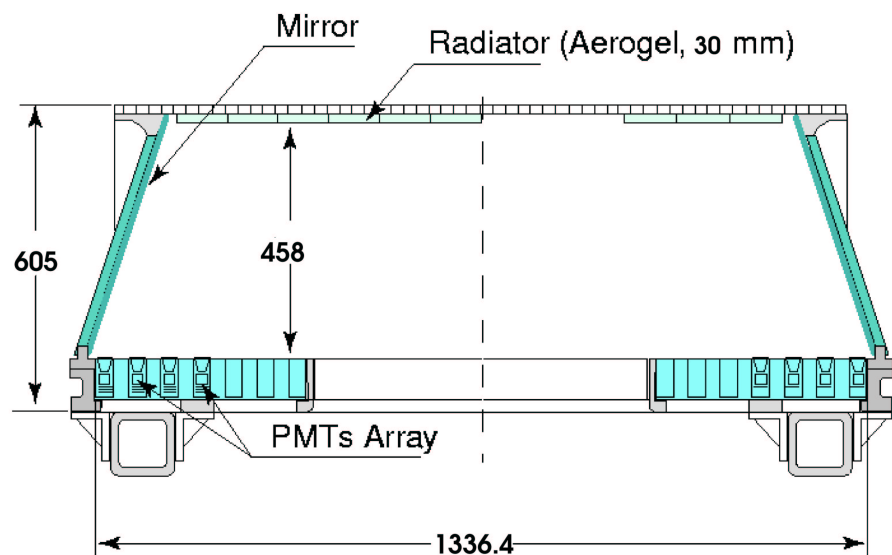
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LIP, Lisbon

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## RICH detector

The Ring Imaging Cerenkov of AMS is a proximity focusing detector with a radiator on the top, a high reflectivity mirror and pixelized photomultiplier tubes.

- velocity measurement  $\frac{\Delta\beta}{\beta} = 0.1\%$
- charge measurement
- redundancy on albedo rejection
- e/p separation

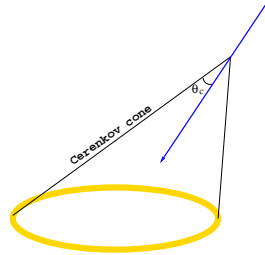


# Cerenkov radiation

## ✓ Cerenkov radiation

a charged particle traveling in a medium with a velocity higher than the light speed radiates photons:

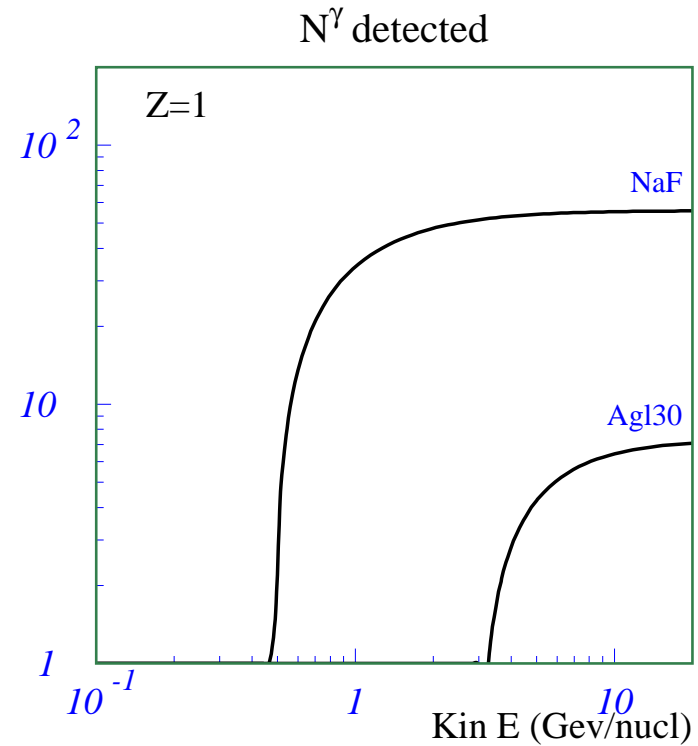
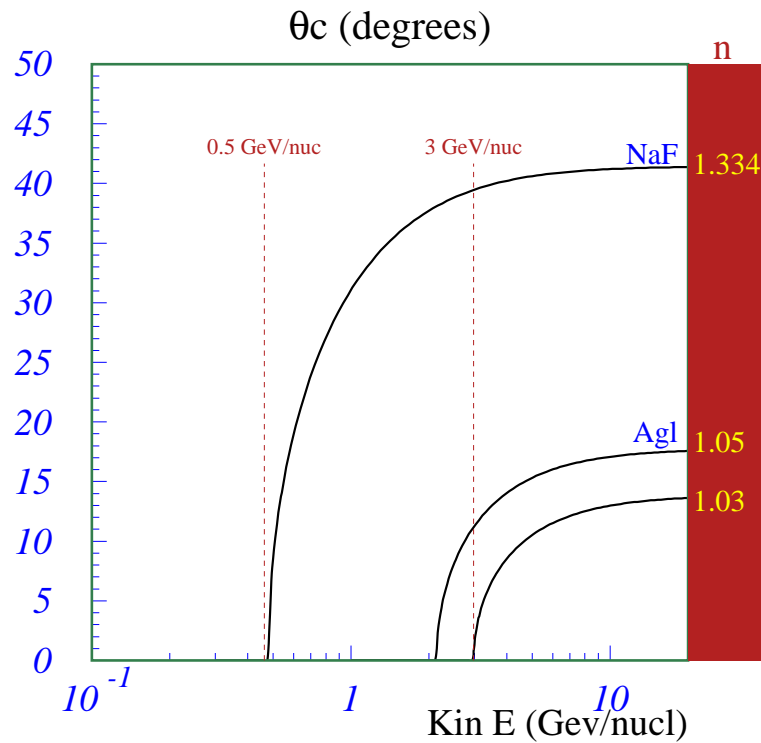
$$\cos\theta_c = \frac{1}{\beta n}$$



## ✓ Light Yield

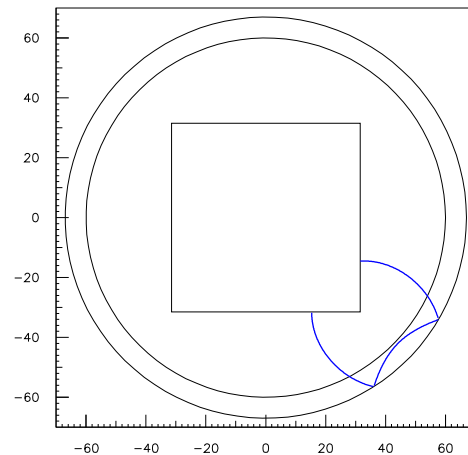
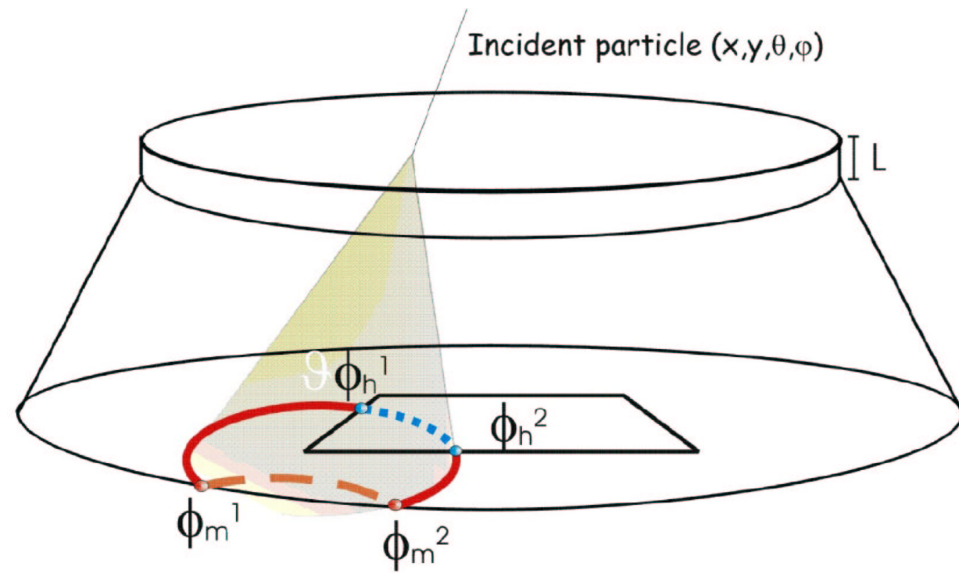
the light yield increases with the radiator thickness (L), the charge (Z), the velocity ( $\beta$ ) and refractive index (n):

$$n_{p.e} \propto Z^2 L \left(1 - \frac{1}{\beta^2 n^2}\right) \int \epsilon dE$$

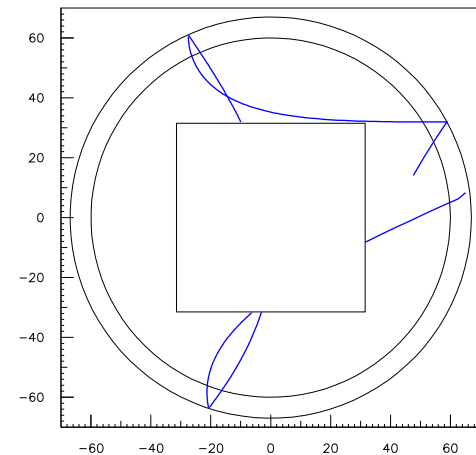


## Fraction of visible photons

- ✓ photons are emitted with an opening angle  $\theta_c$  and with an azimuthal angle  $\varphi$  ranging from  $(0, \pi)$
  - ✓ a fraction of these photons are *lost* due to (*geometrical acceptance*):
    - ⇒ escaping from radiator (lateral walls)
    - ⇒ total reflecting medium transitions
    - ⇒ falling into non-active readout area
- $\sim 63 \times 63 \text{ cm}$



aerogel ( $n=1.030$ )

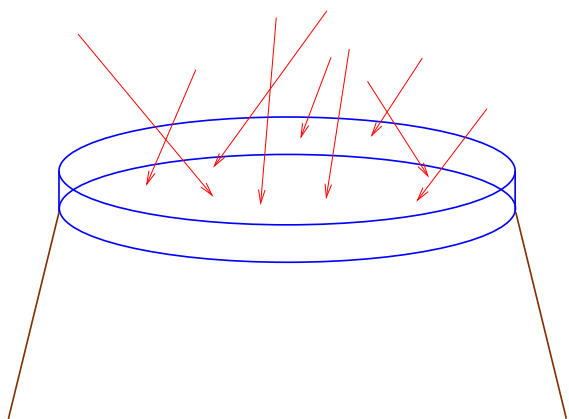


NaF

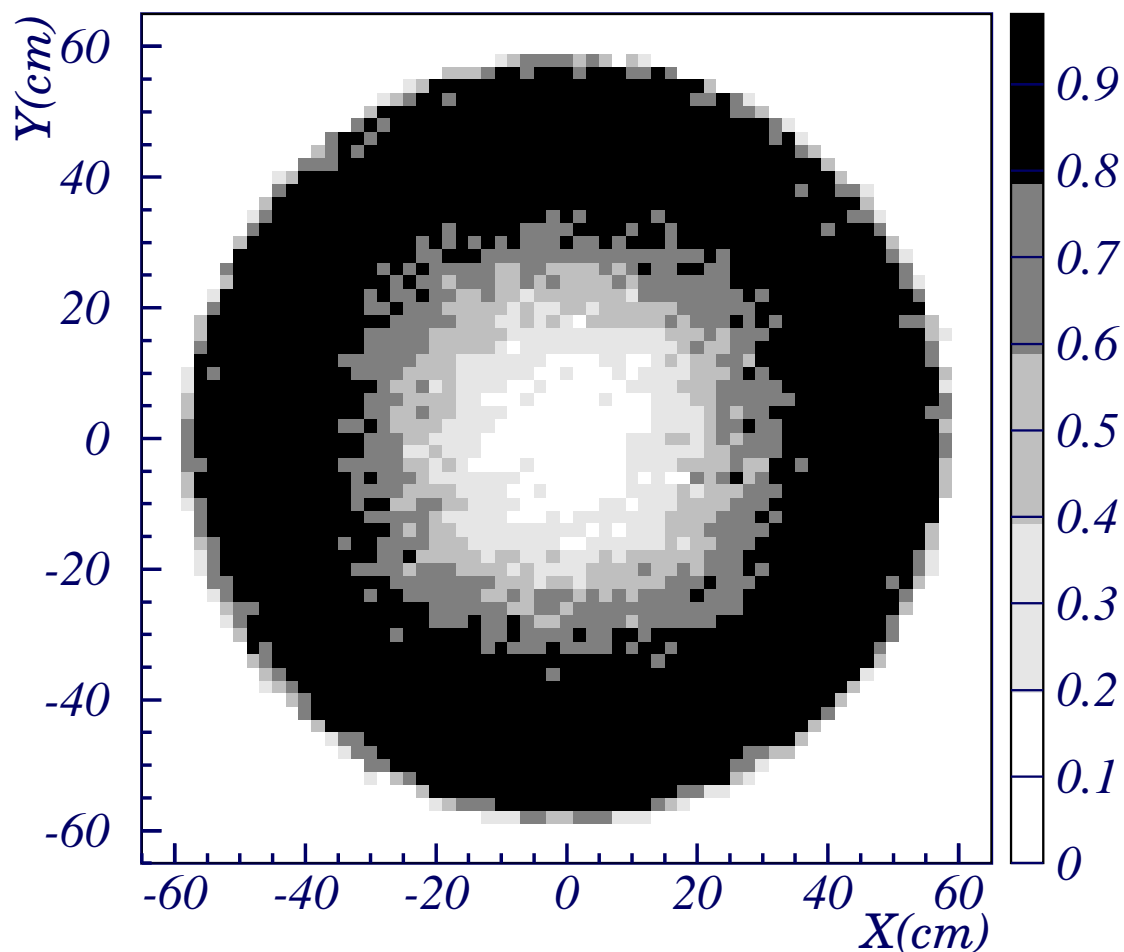
### typical patterns

- ✓ for aerogel ( $n=1.030$ )
- ✓ for NaF ( $n=1.334$ )

## Event geometrical acceptances with Aerogel

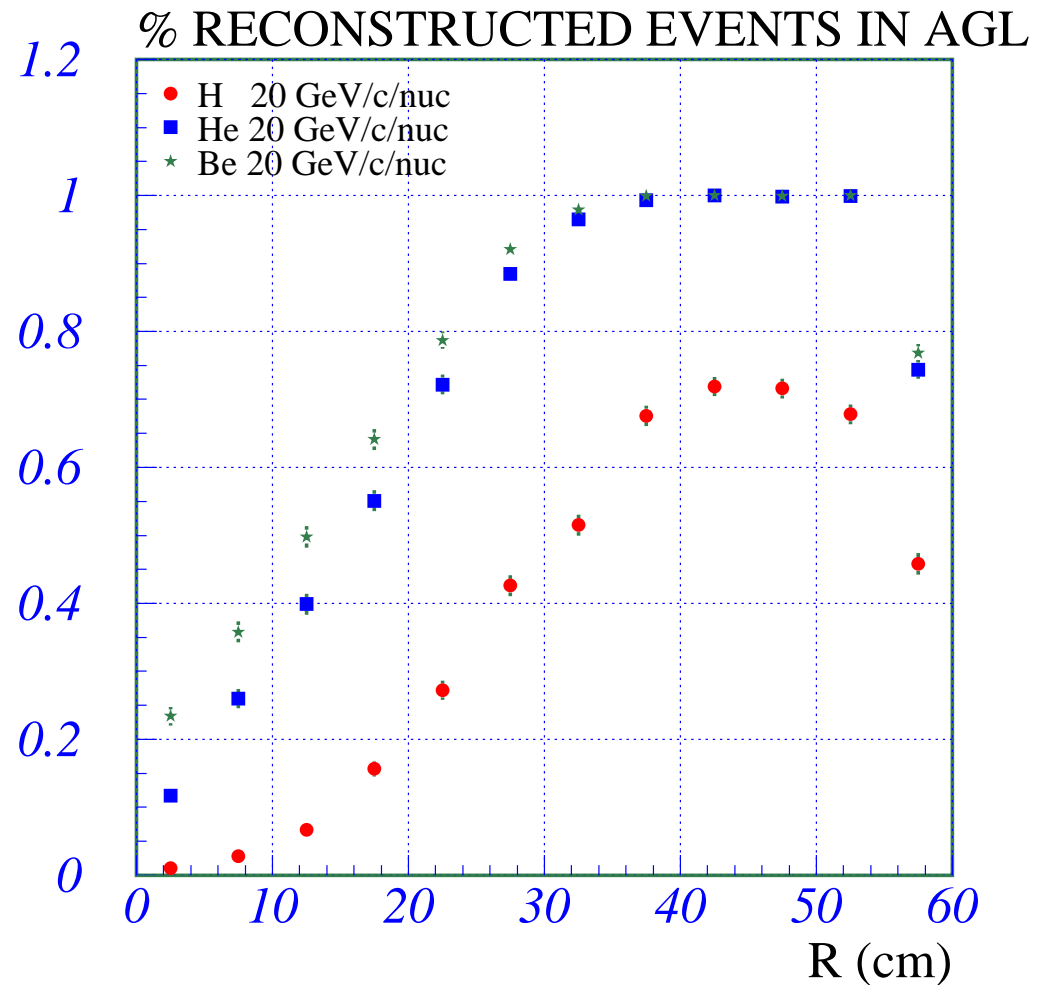
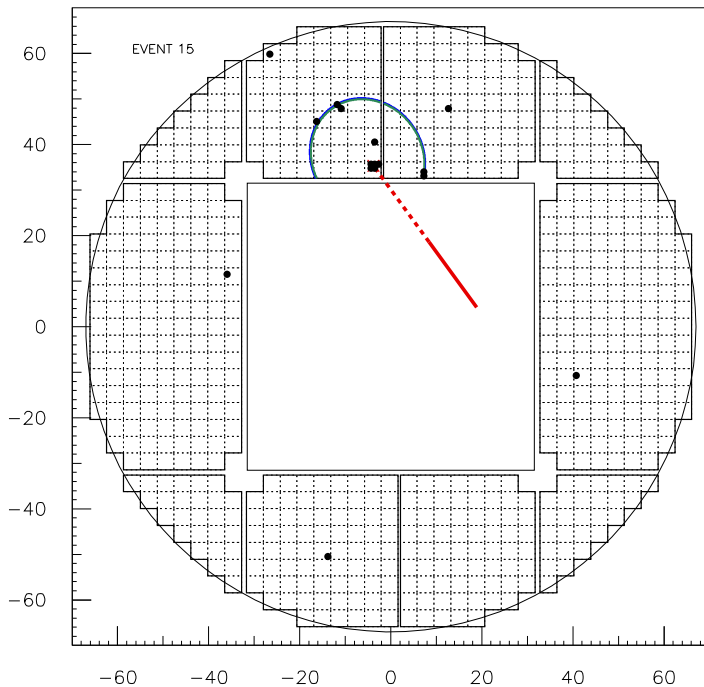


- ✓ particles are impinging within the AMS acceptance on top of aerogel radiator ( $n=1.03$ )
- ✓ the fraction of detected photons depends on:
  - particle impact point on the radiator
  - particle direction



## Aerogel: event reconstruction efficiencies

- ✓ cerenkov angle is reconstructed for **protons**, **heliums** and **berylliums**
  - ⇒  $\geq 3$  hits required
- ✓ reconstruction efficiency drops as far as particles get closer to the radiator center
  - ⇒ large fraction of photons lost on *hole*



# The Cerenkov signal and the $\theta_c$ resolution

Cerenkov signal depends on:

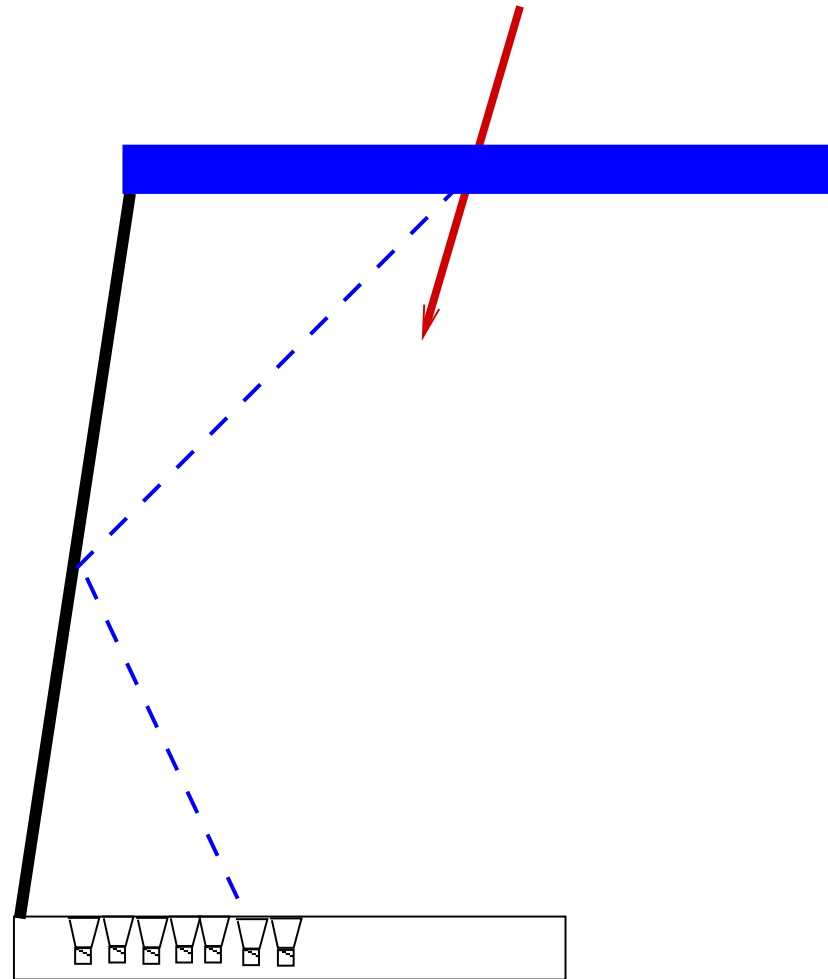
- ⇒ radiator:
  - ✓ refractive index (n)
  - ✓ chromaticity
  - ✓ interactions (scattering, absorption)
- ⇒ geometrical acceptance
- ⇒ light guide
- ⇒ pmt

Resolution:

$$\frac{\Delta\beta}{\beta} = \tan \theta_c \Delta\theta_c / \sqrt{N^{hits}}$$

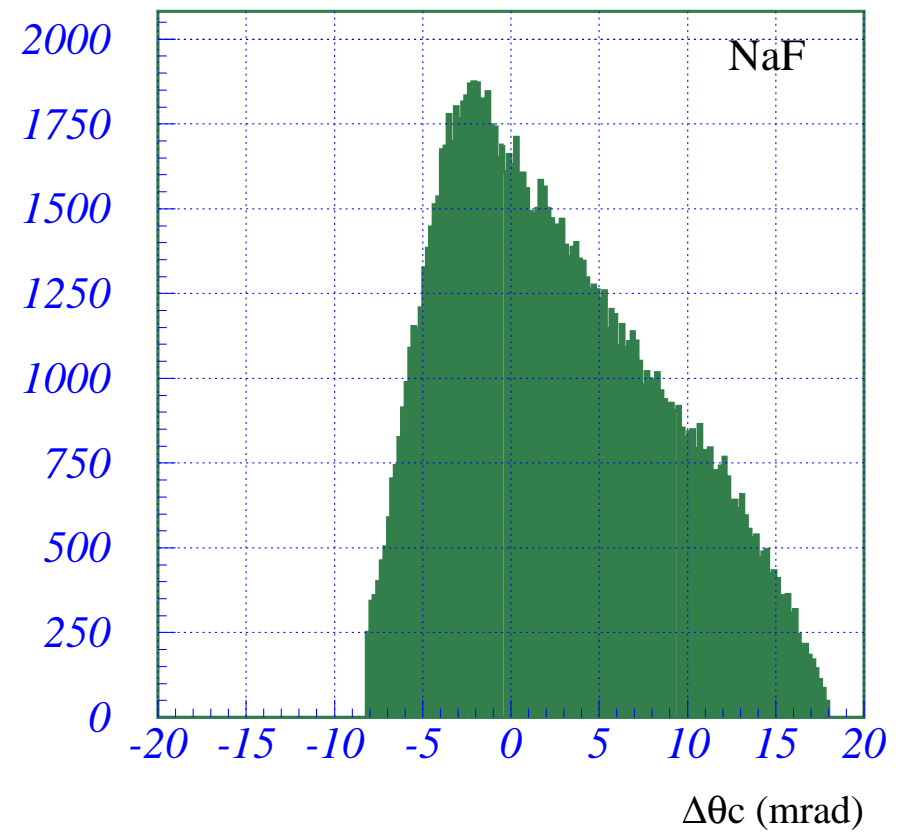
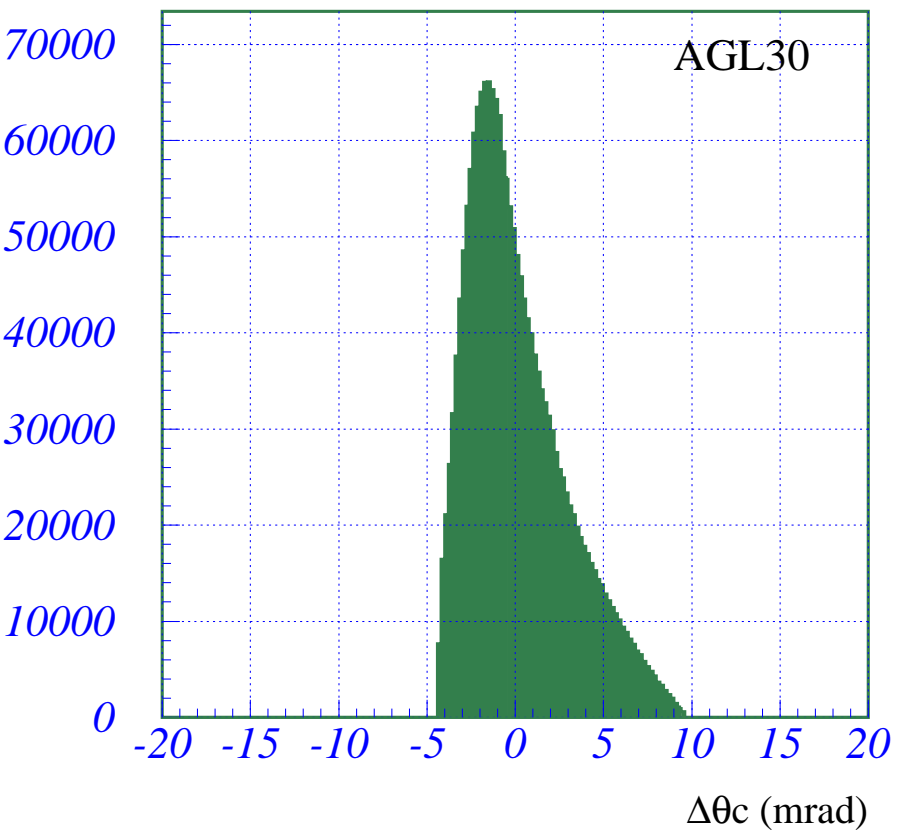
the  $\theta_c$  uncertainty deals with:

- ⇒ pixel size (8.5 mm)
- ⇒ chromaticity
- ⇒ radiator thickness



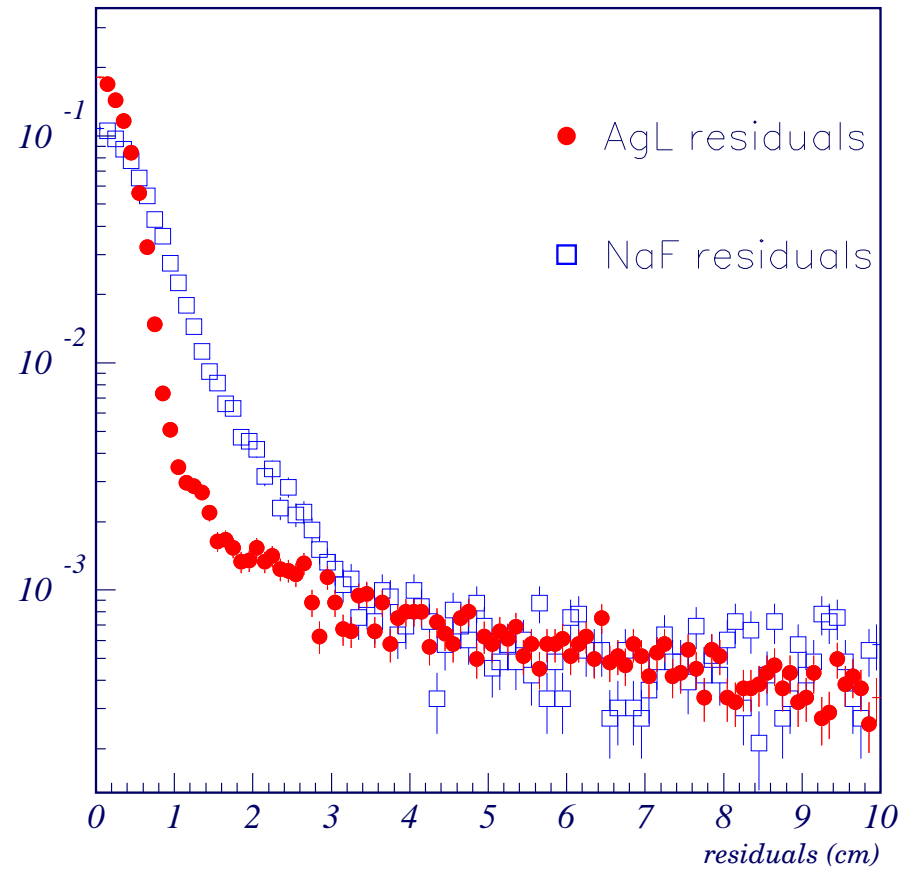
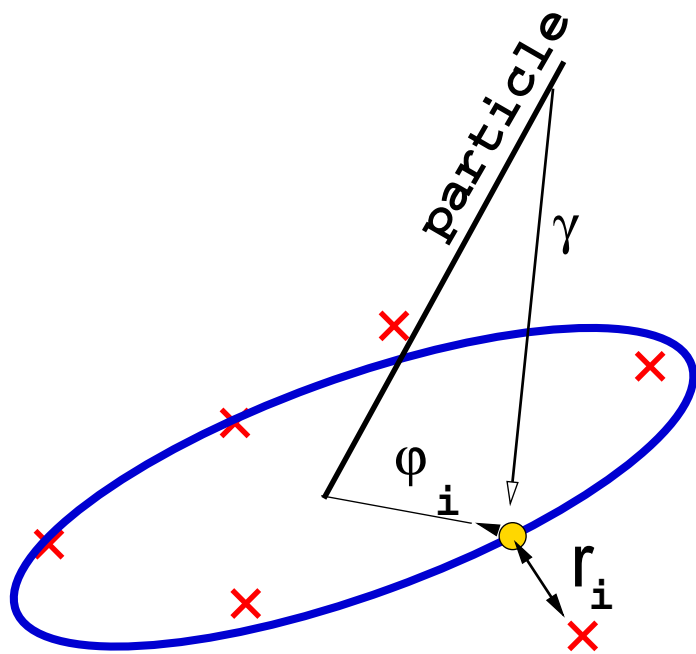
## NaF.vs.Agl: chromaticity dispersion

$$\cos\theta_c(\lambda) = \frac{1}{\beta n(\lambda)}$$



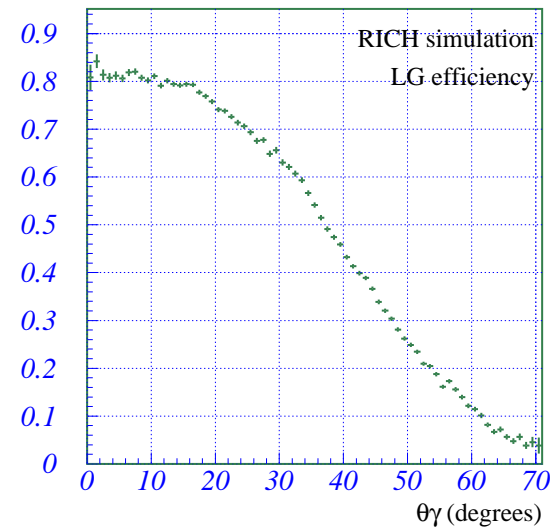
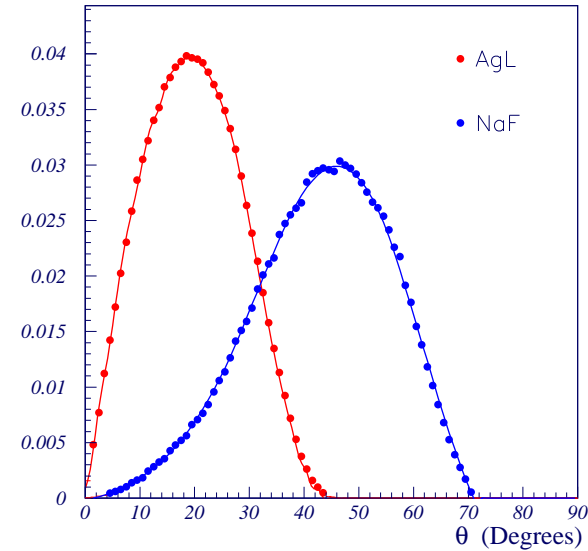
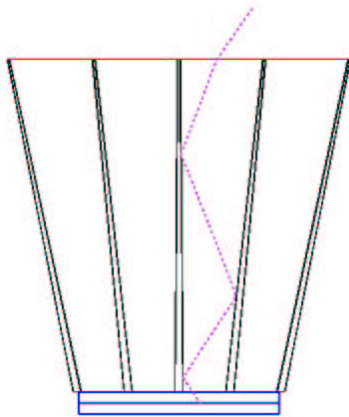


# NaF.vs.Agl: hit residuals



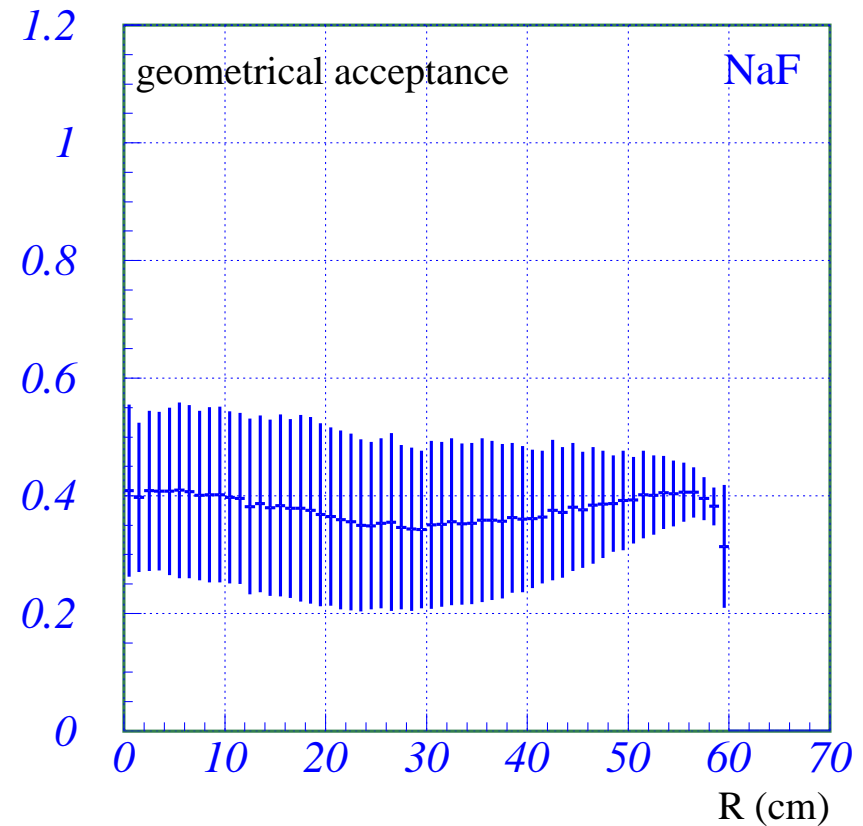
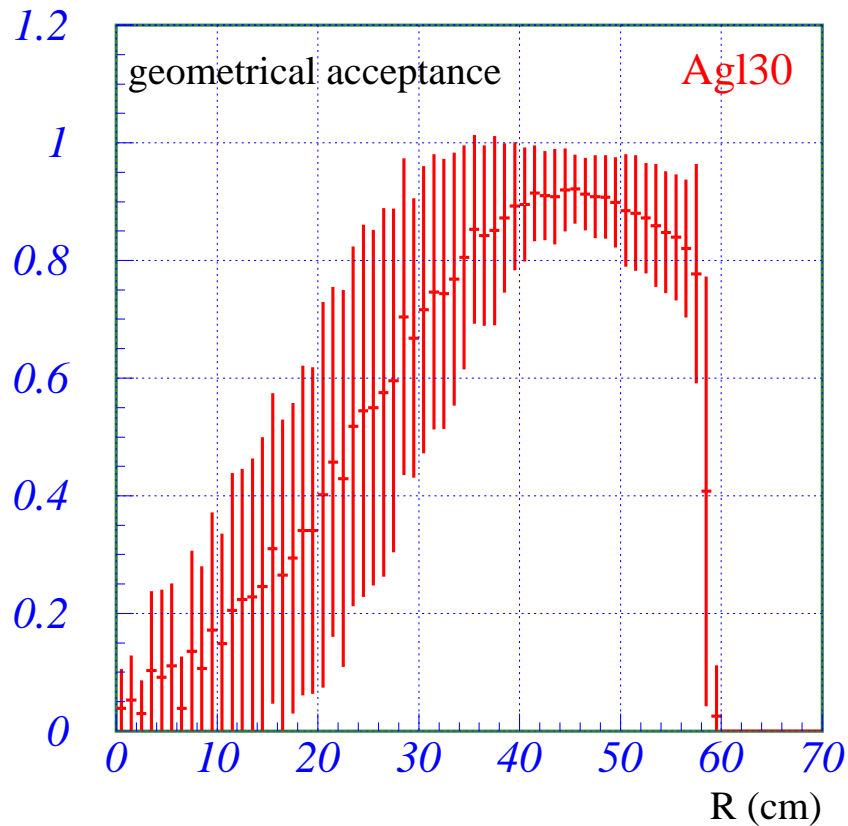
# NaF.vs.Agl: Light Guide Efficiency

- ⇒ The light guide efficiency depends on the photon entrance angle ( $\theta_\gamma$ )
- NaF radiated photons have larger entrance angles and therefore lower efficiencies



## NaF vs. AgI: geometrical acceptance

- Fraction of lost photons as function of the particle impact point distance to the radiator center
- essentially due to the non-active region on PMT matrix

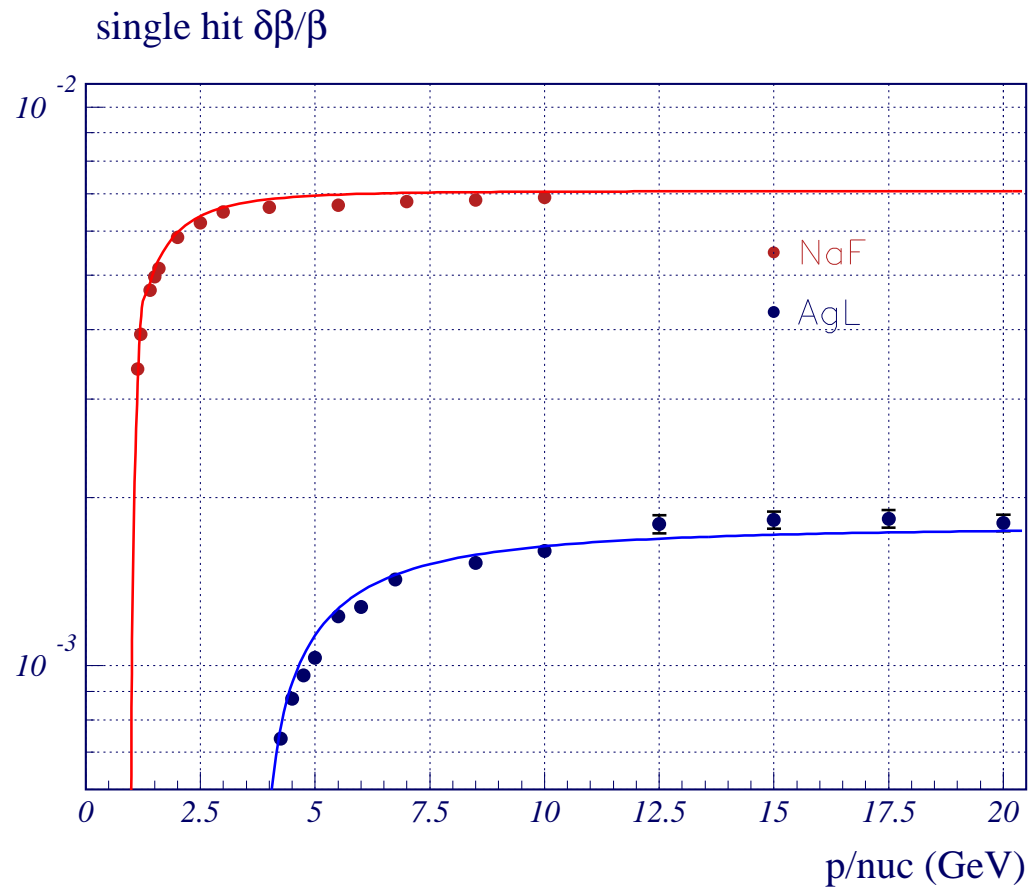


## NaF.vs.Agl

### Confronting Aerogel and Naf

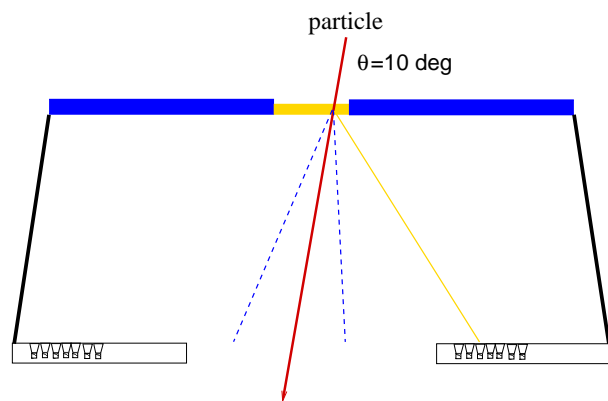
Effect on the number of cerenkov hits used on $\theta_c$ reconstruction	Agl	NaF
<input type="checkbox"/> nb of radiated photons (/cm)	➤	➤
<input type="checkbox"/> radiator chromaticity	➤	➤
<input type="checkbox"/> radiator interactions	➤	➤
<input type="checkbox"/> light guide efficiency	➤	➤
<input type="checkbox"/> geometrical acceptance ( $R < 15 - 20 \text{ cm}$ )	➤	➤
<input type="checkbox"/> geometrical acceptance ( $R > 20 \text{ cm}$ )	➤	➤

# NaF.vs.Agl: resolutions

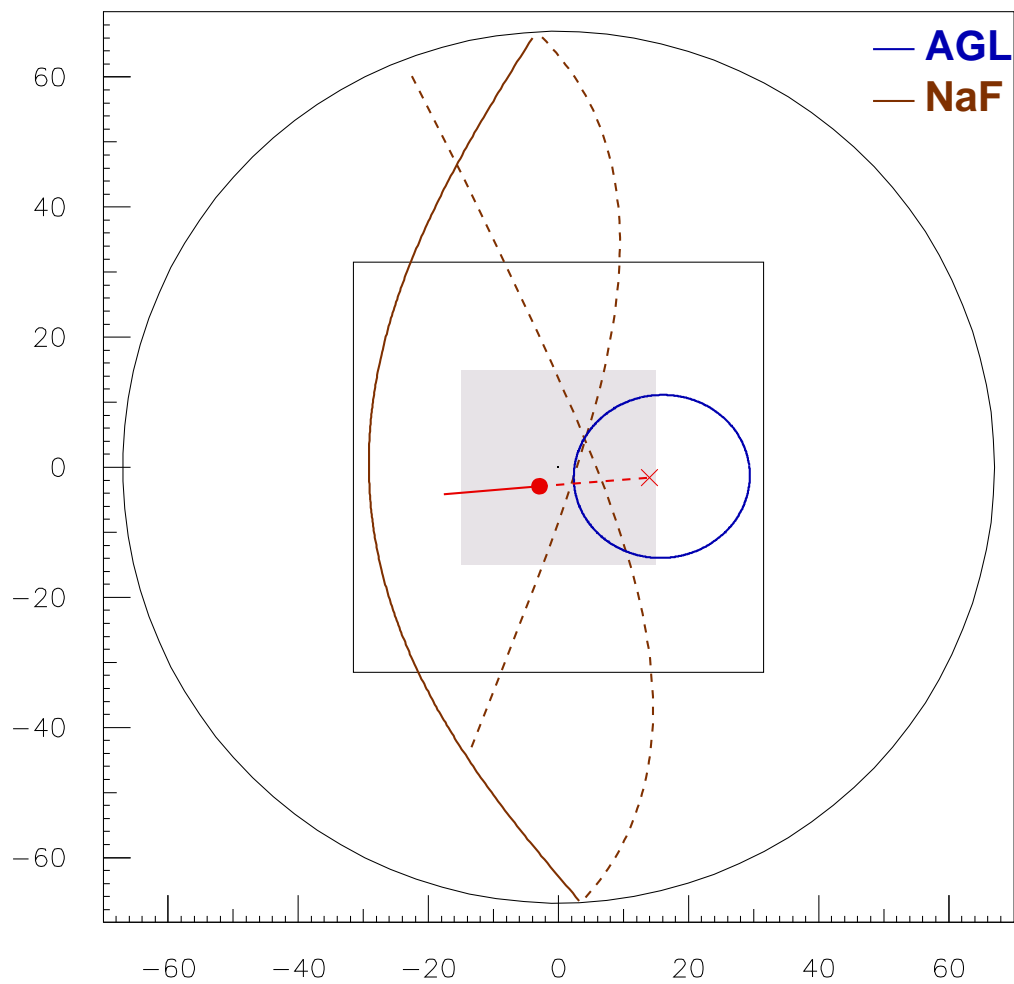


## Solving the problem..

- A larger refractive index radiator in the central region

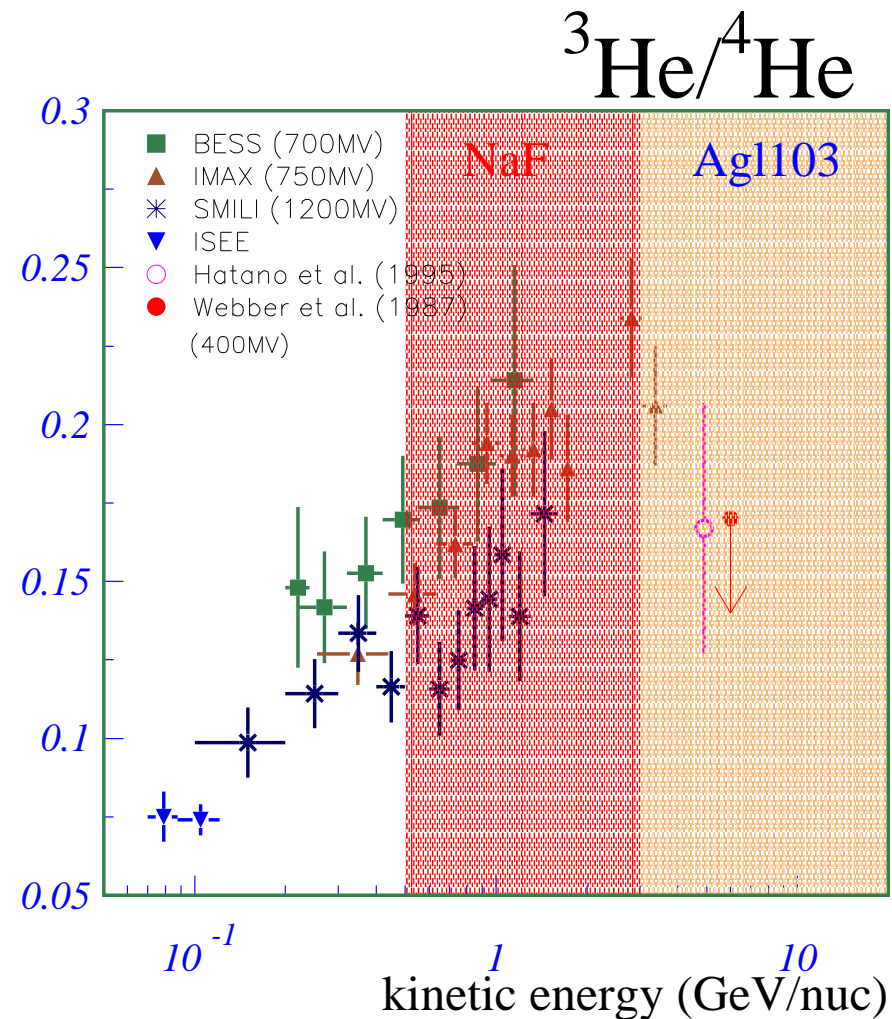


- photons emitted at larger cerenkov angles
- more photons detected in the active area



## Physics arguments: helium isotopes

- ⇒ The propagation history of the helium can be probed by measuring the isotopic ratio  ${}^3\text{He}/{}^4\text{He}$
- ⇒  ${}^3\text{He}$  is essentially secondary and comes from the spallation of  ${}^4\text{He}$
- ⇒ aerogel in AMS will provide isotopic ratios from  $E_{kin} \simeq 3 \text{ GeV/nuc}$
- ⇒ the integration of NaF in the Rich radiator will allow to measure isotopic ratios down to  $E_{kin} \simeq 0.5 \text{ GeV/nuc}$



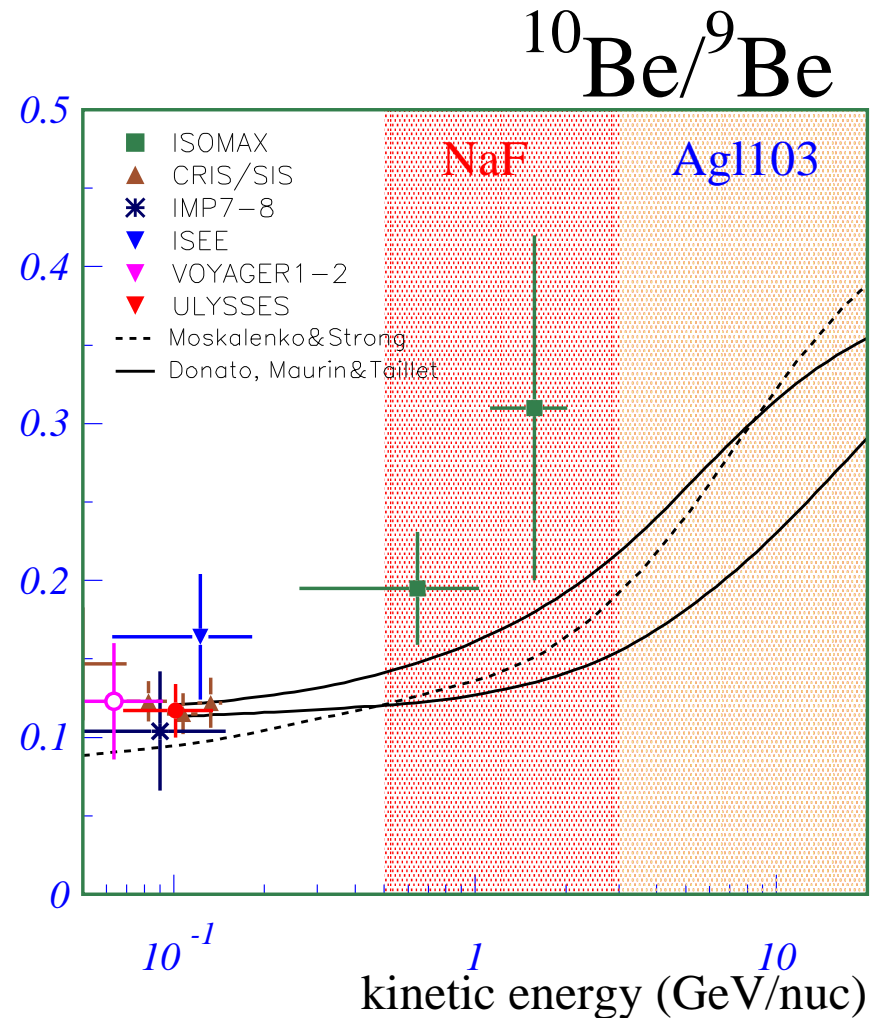
## Physics arguments: beryllium

- Measurement of ratio  $^{10}\text{Be}/^9\text{Be}$  give us information about **confinement of cosmic rays** in the Galactic volume and is sensitive to **different propagation models**

$t_{1/2}(^{10}\text{Be}) \sim 1.5 \times 10^6 \text{ yrs}$

- aerogel in AMS will provide isotopic ratios from  $E_{kin} \simeq 3 \text{ GeV/nuc}$

- the integration of NaF in the RICH radiator will allow to measure isotopic ratios down to  $E_{kin} \simeq 0.5 \text{ GeV/nuc}$



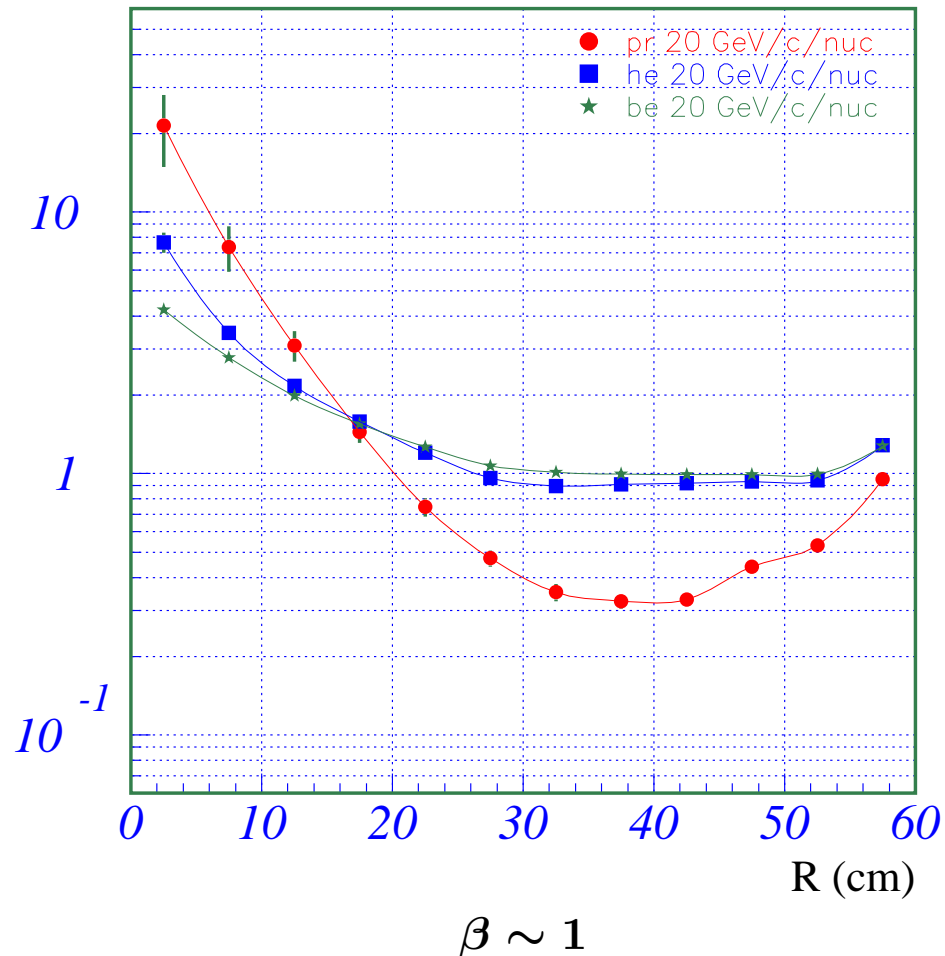


## Which size for NaF?

- ⇒ A minimal number of 3 hits required on  $\theta_c$  reconstruction
- ⇒ The ratio of the events reconstructed on **NaF** and **Agl30** as function of the incident particle distance to the radiator center

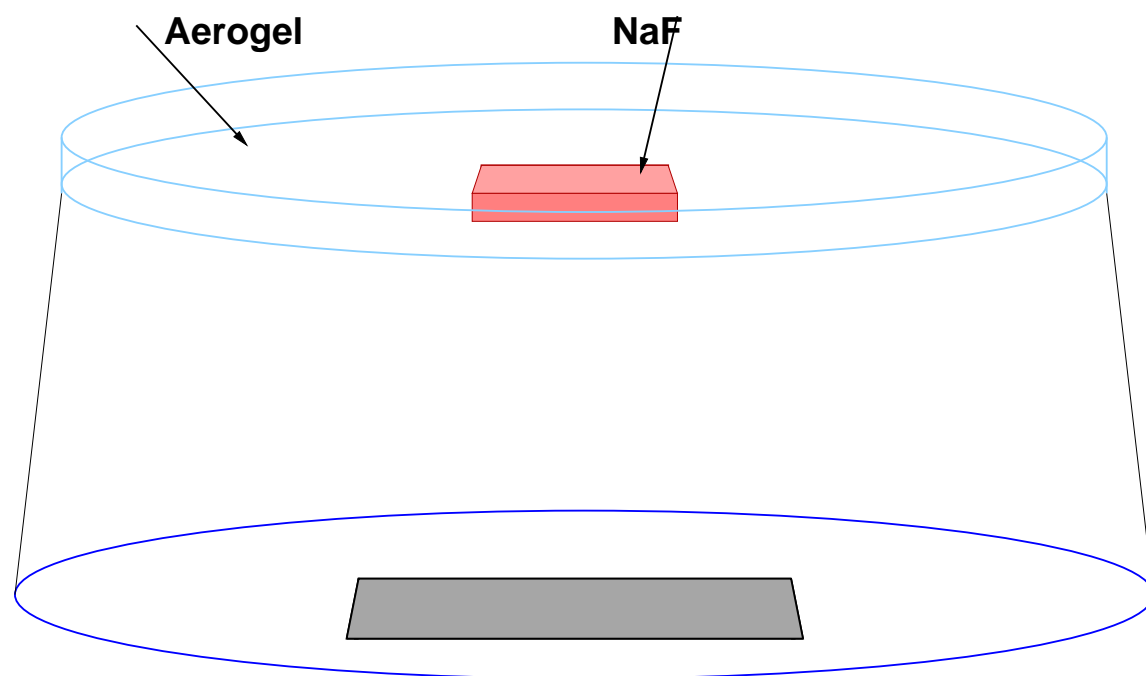
$$\frac{\% \text{ Reevents}(\text{NaF})}{\% \text{ Reevents}(\text{Agl30})}$$

- ⇒ NaF reconstruction is dominant for  $R$  lower than  $\sim 20 \text{ cm}$

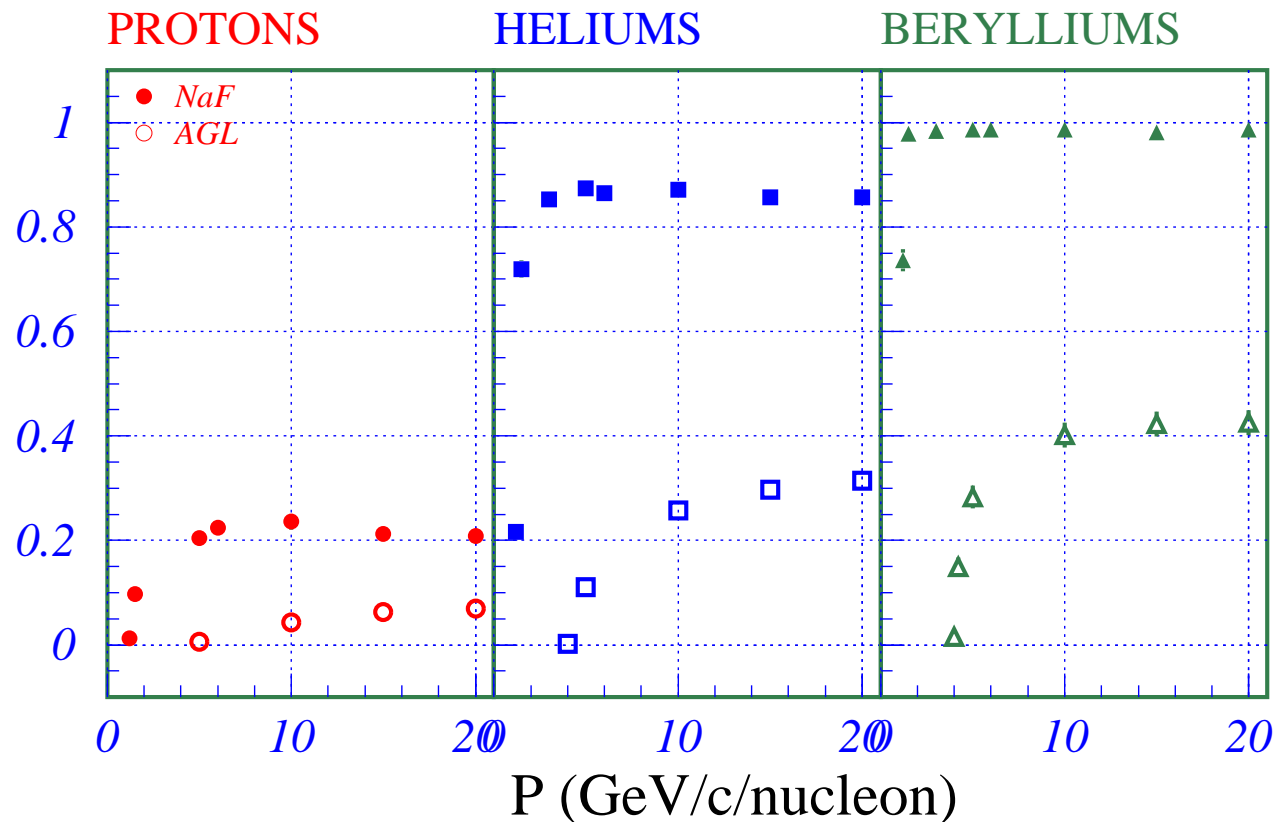


## Dual radiator configuration

- ⇒ a square of sodium fluoride (NaF) with  $\sim 30 \times 30 \times 0.5 \text{ cm}^3$  placed in the center of of the RICH radiator
- ⇒ it covers about 10% of the RICH acceptance
- ⇒ an amount of matter corresponding to  $\sim 4\%$  of  $X_0$  (*aerogel is  $\sim 3\%$* )
- ⇒ important implications:
  - ✓ RICH acceptance increases which implies larger reconstruction efficiencies
  - ✓ kinetic energy range is extended down to values around  $0.5 \text{ GeV/nuc}$



## NaF.vs.Agl: reconstruction efficiency (30 cm square)

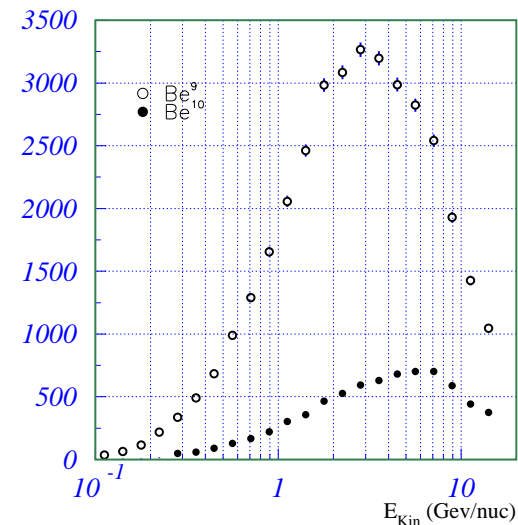
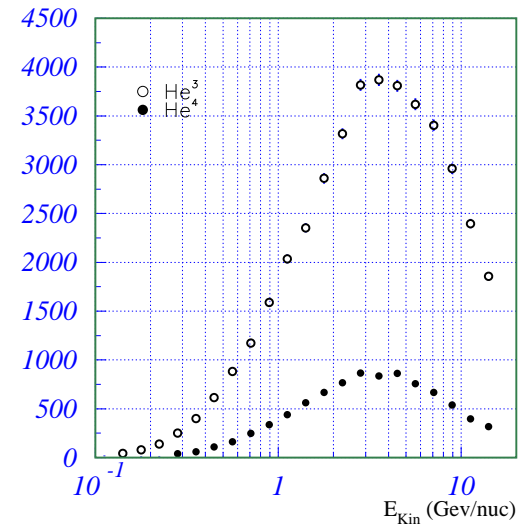


⇒ The fraction of particles impinging on the NaF area (30 cm square) and being reconstructed ( $N_{hits} > 2$ ) depends strongly on the charge

*NaF efficiency reaches 100% for Berylliums*

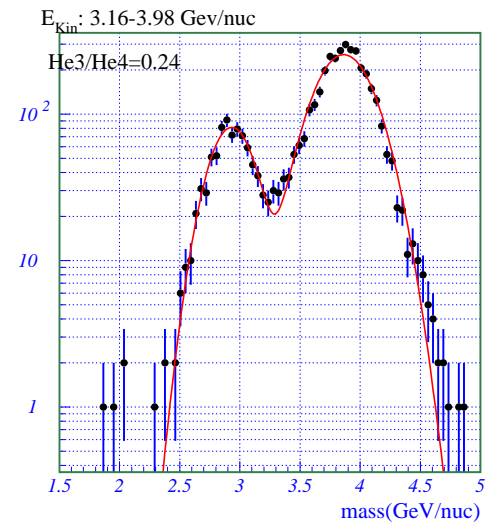
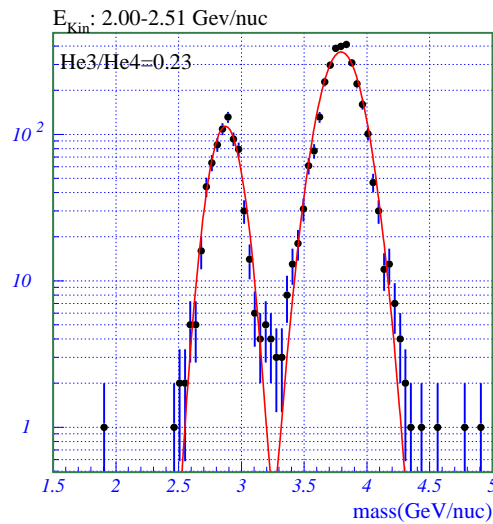
## Simulation of helium and beryllium nuclei

- ⇒ A statistics of  ${}^3\text{He} \sim 85 \times 10^3$  and  ${}^4\text{He} \sim 421 \times 10^3$  nuclei, were simulated through the RICH  
6 hours of data taking
- ⇒ A statistics of  ${}^{10}\text{Be} \sim 75 \times 10^3$  and  ${}^9\text{Be} \sim 349 \times 10^3$  nuclei, were simulated through the RICH  
6 months of data taking
- ⇒ geomagnetic field taken into account  
modulation of the nuclei energy with the ISS location
- ⇒ momentum uncertainty folded

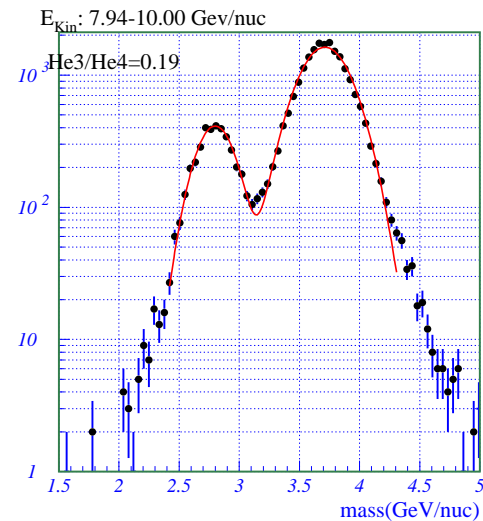
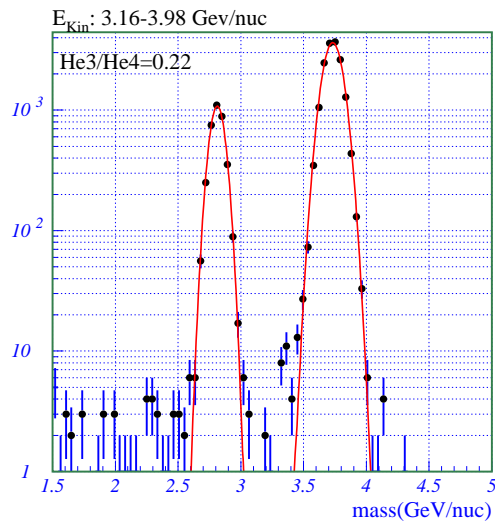


# Helium isotopic separation

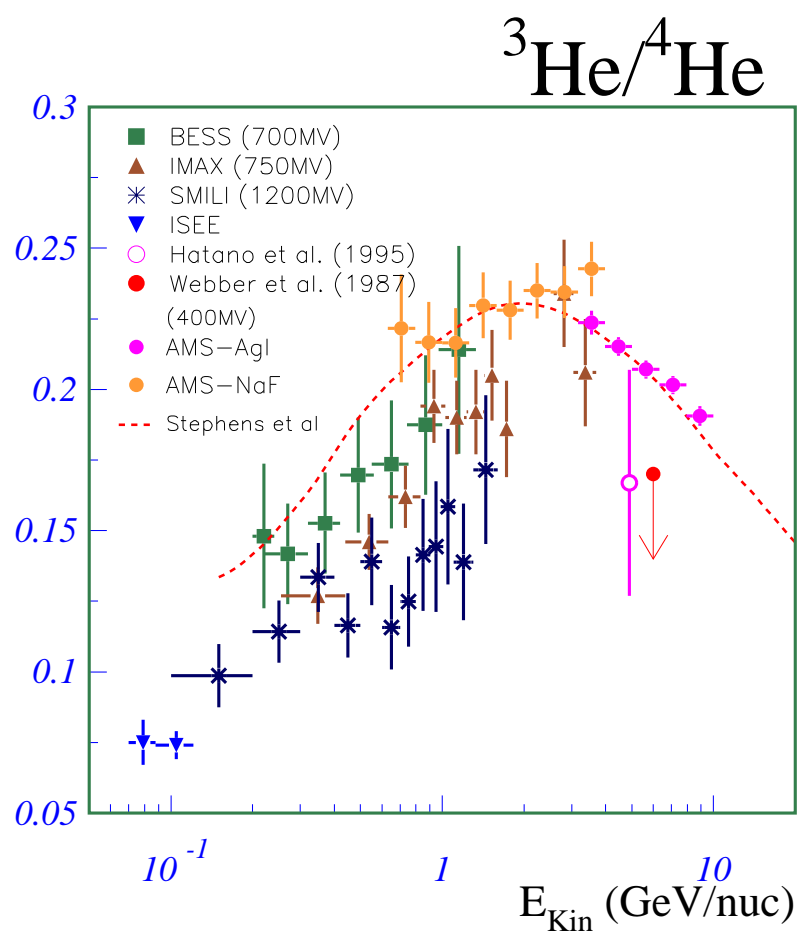
NaF



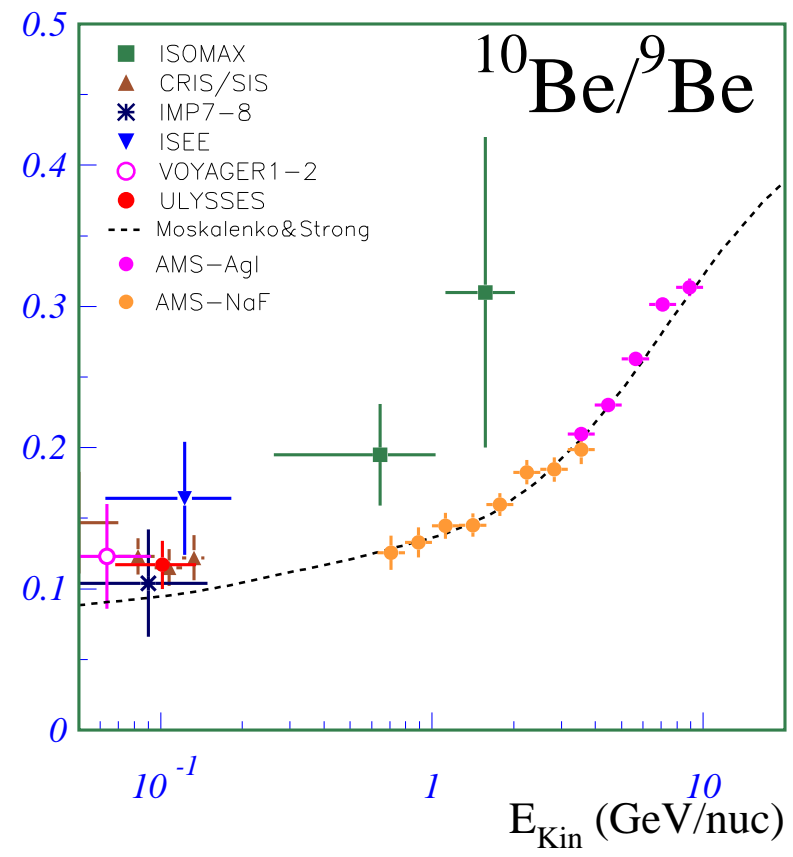
Ag130



# Reconstructed isotopic ratios for He and Be



**AMS data: 6 hours**



**AMS data: 6 months**

## Conclusions

- ⇒ The possibility of having a mixed radiator configuration with both a large and a low refractive index radiators, was studied
- ⇒ Aerogel radiator shows low event geometrical acceptances for particles impinging close to the radiator center
- ⇒ The placement of a NaF radiator at the center of the radiator plane ( $30 \times 30 \text{ cm}^2$ ) increases substantially the number of reconstructed events ( $N_{\text{hits}} > 2$ ), when compared with aerogel
- ⇒ At last (but not at least) the introduction of a NaF radiator allows AMS to cover the complete spectrum of helium and beryllium isotopic measurements from  $0.1 \text{ GeV}/\text{nuc}$  up to around  $10 \text{ GeV}/\text{nuc}$